LONG-PERIOD TRANSITING EXOPLANETS

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The "population" of exoplanets

Transiting exoplanets











Kepler









NGC6791

Credit NASA





Who cares?

The population of exoplanets





Burke et al. (2015)





Burke et al. (2015)



The frequency of Solar System analogs









Why Kepler?

Why Kepler?



Why Kepler?



Ingredients for population inference



Today's punch line





Candidates (green) from DFM et al. (in prep); Data from NASA Exoplanet Archive

How to find a transiting planet...

short-period How to find a transiting planet...





The anatomy of a transit





... just do the inference?
no.



Filter the data to "remove" systematics



Template-based grid of likelihoods

(restricted to systems with >2 transits)



Remove false alarms by "visual inspection"



Filter the data to "remove" systematics



Template-based grid of likelihoods

(restricted to systems with >2 transits)



Remove false alarms using magic



~190,000 target stars



Template-based grid of likelihoods

(restricted to systems with >2 transits)



Remove false alarms using magic









e.g. Wang et al. (2015); Uehara et al. (2016); Kipping et al. (2016)

Ingredients for population inference



Ingredients for population inference



A fully-automated detection method

1

Filter the data to "remove" systematics



Template-based grid of likelihoods

(restricted to high signal-to-noise candidates)







* fast GP regression at github.com/**dfm/george**

Still too expensive...

...use BIC*

* don't take this slide out of context



* github.com/dfm/george; Ambikasaran, DFM et al. (2014)





DFM et al. (in prep)

Why not Machine Learning?

(e.g. supervised classification)

The Kepler data are not Big[™].

all but ~0.02% of

The Kepler data are Boring[™].*

Results



DFM et al. (in prep)

1

Filter the data to "remove" systematics



Template-based grid of likelihoods

(restricted to high signal-to-noise candidates)





~40,000 target stars



Template-based grid of likelihoods

(restricted to high signal-to-noise candidates)





3



* some contamination from EB secondary eclipses



Candidates (green) from DFM et al. (in prep); Data from NASA Exoplanet Archive



DFM et al. (in prep)

Ingredients for population inference





DFM et al. (in prep)



20 0

30 60

occurrence rate in period range 2 – 25 years $\begin{array}{ll} R_E - R_N & R_N - R_J \\ \sim 0.40 & \sim 0.17 \end{array}$

per G/K- dwarf, per In-radius, per In-period

DFM et al. (in prep) compare with Bryan et al. (2016); Shvartzvald et al. (2016)

Summary



Fully automated discovery of longperiod transiting exoplanets in Kepler archival data



Empirical measurement of search completeness



Estimate of the occurrence rate of long-period exoplanets

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