

Reconstructing cosmic reionization

Layne Price

Carnegie Mellon University

Hydrogen reionization

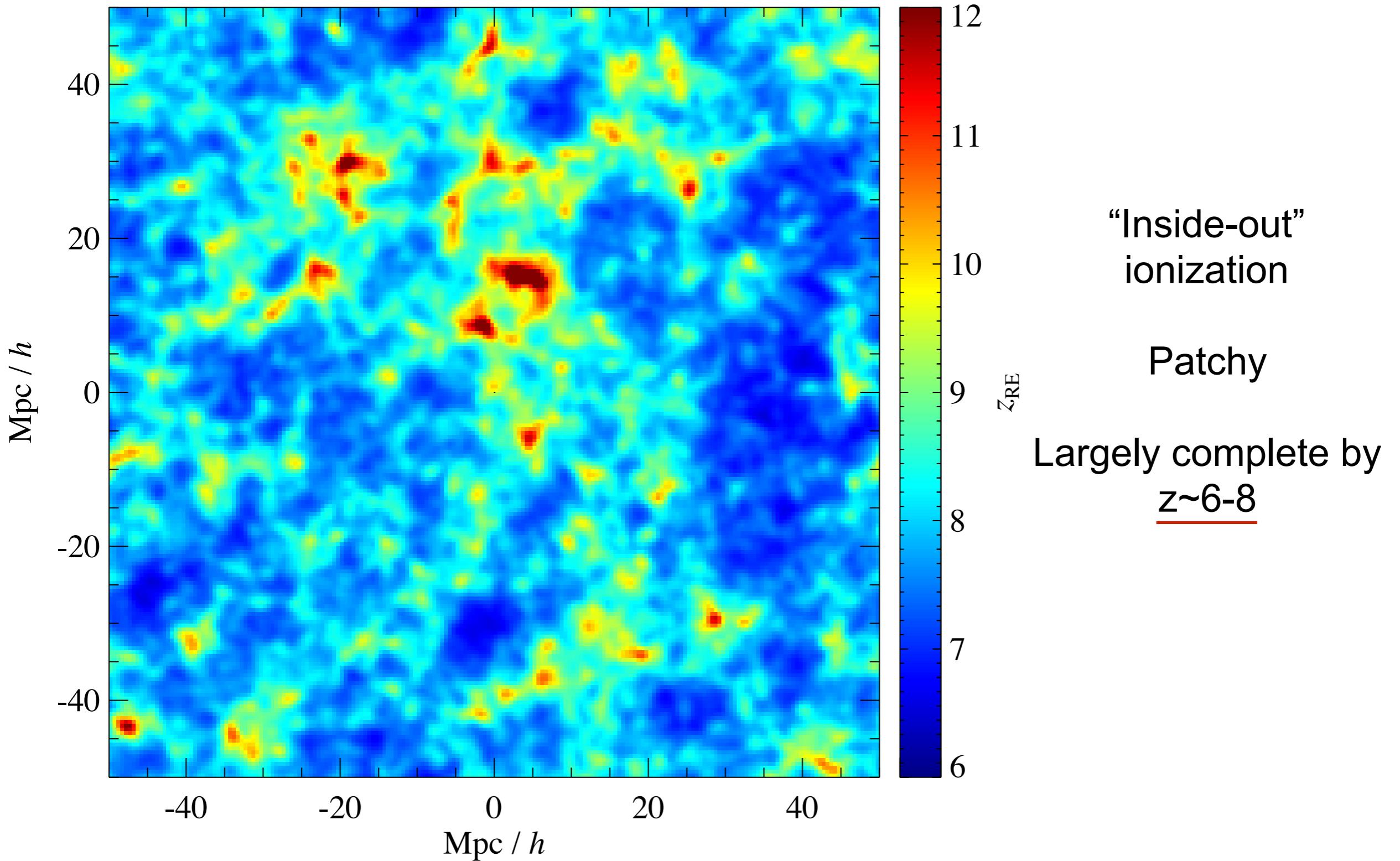
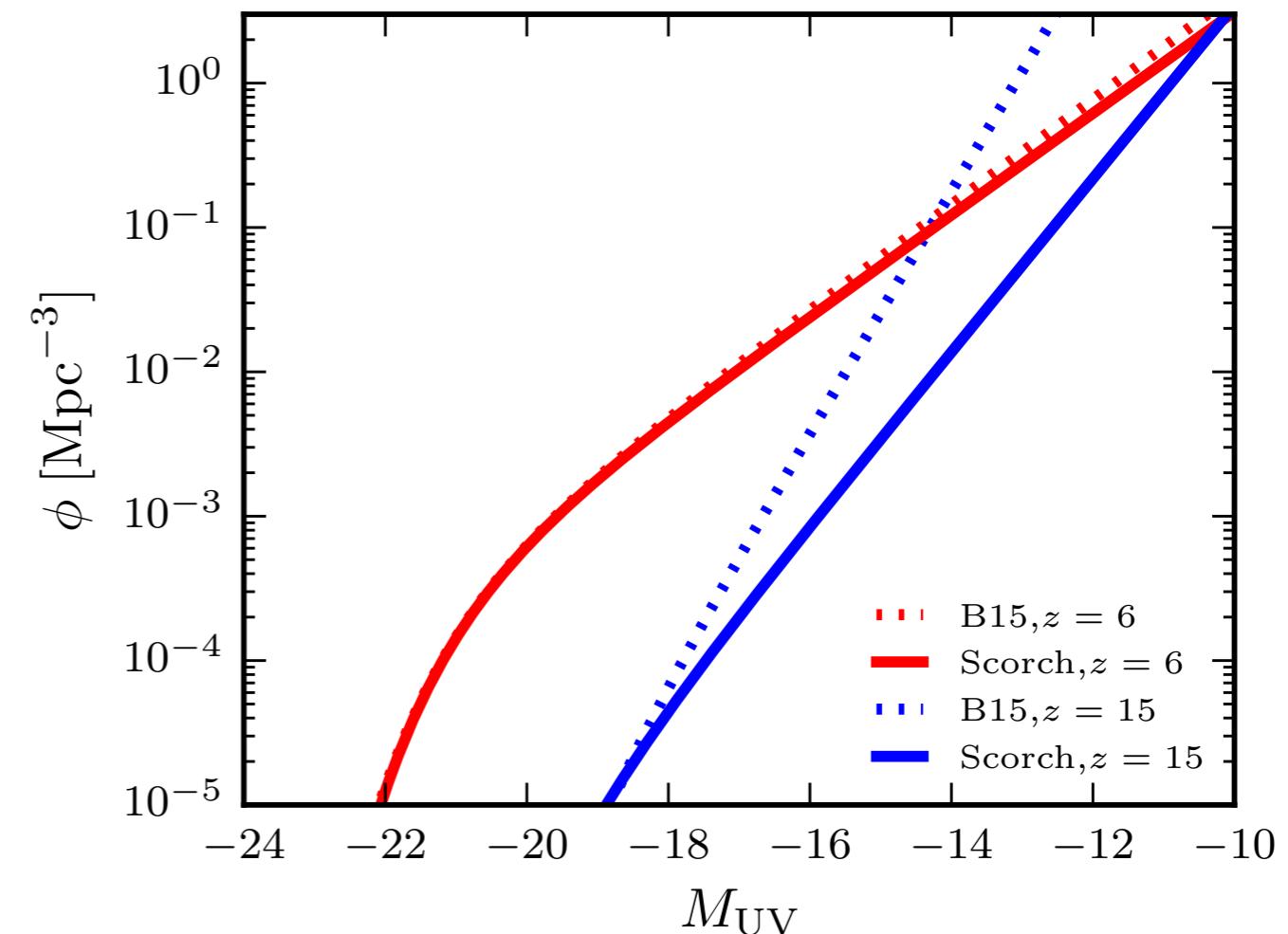
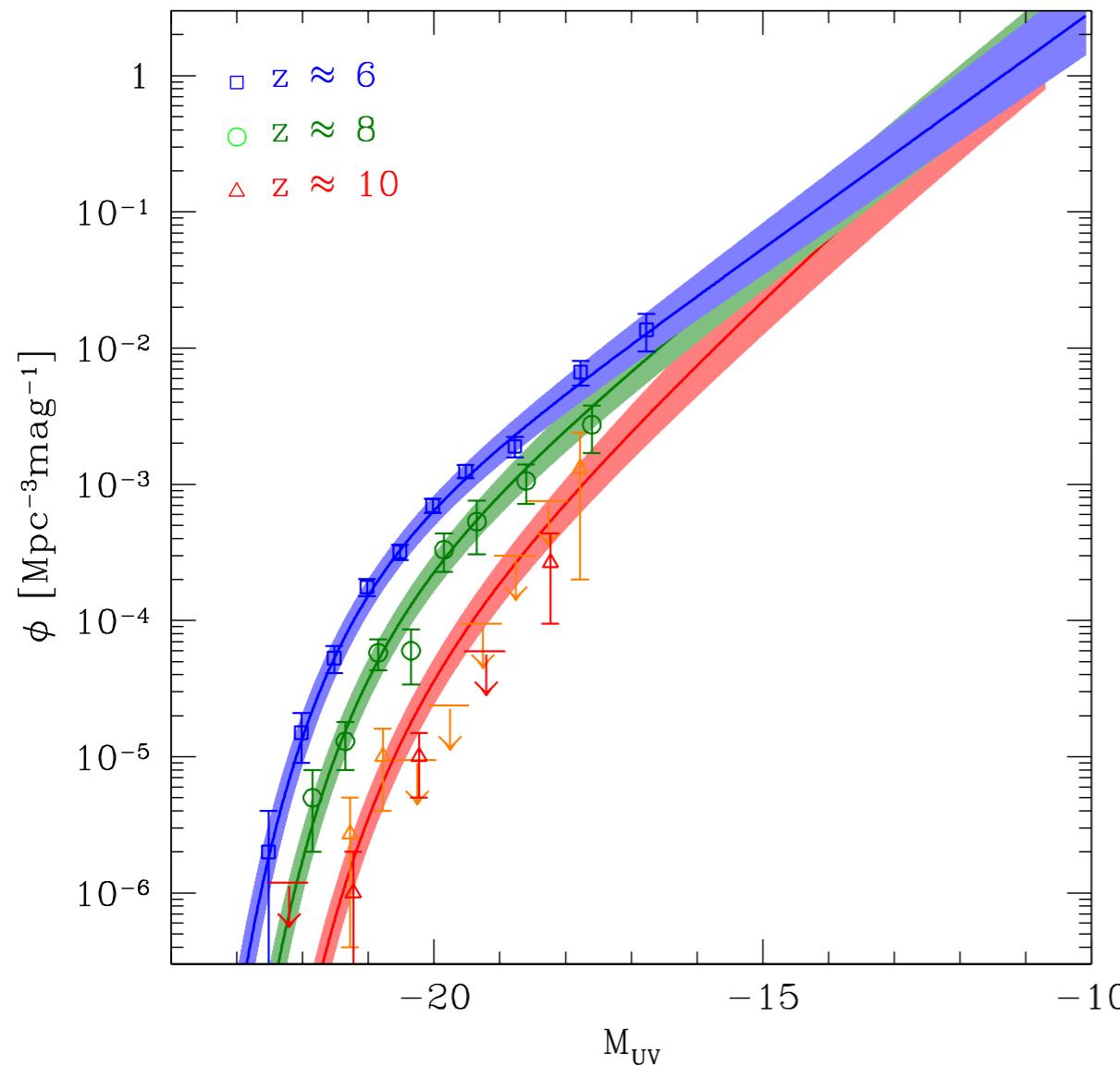


Fig. from Battaglia, et al., ApJ, 1211.2821

The role of low-luminosity galaxies



Extrapolating Schechter fit to HST data to
lower luminosity and higher redshift

Escape of ionizing photons into the IGM

Number of galaxies at given UV luminosity:

$$\phi_* \left(\frac{L}{L_*} \right)^\alpha \exp \left(\frac{L}{L_*} \right)$$

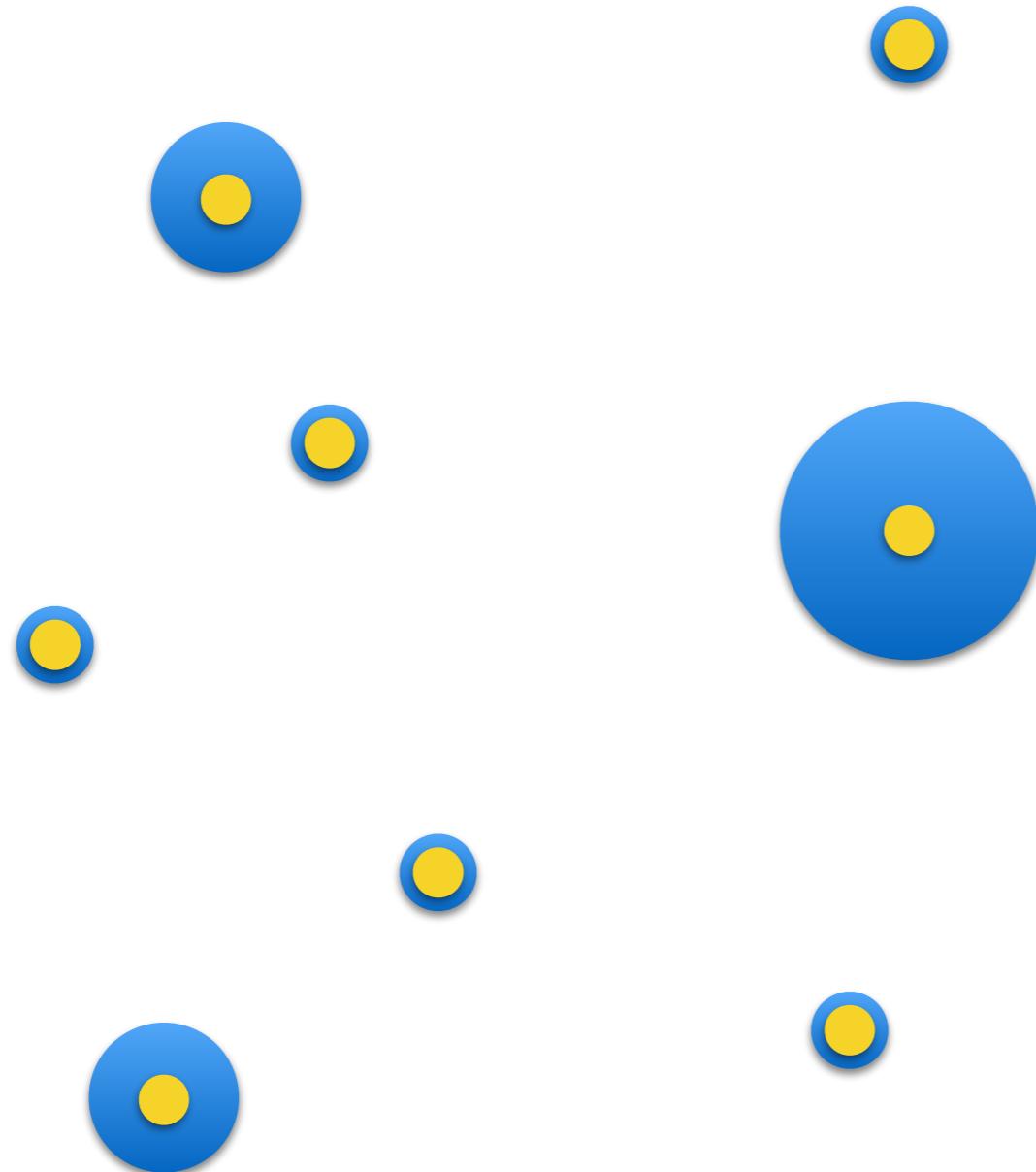
Converting UV luminosity to H ionizing luminosity:

$$L_{\text{ion}} = \xi_{\text{ion}} L_{\text{UV}}$$

Fraction of intrinsic flux that escapes into IGM:

$$f_{\text{esc}}(z)$$

Modeling reionization bubbles

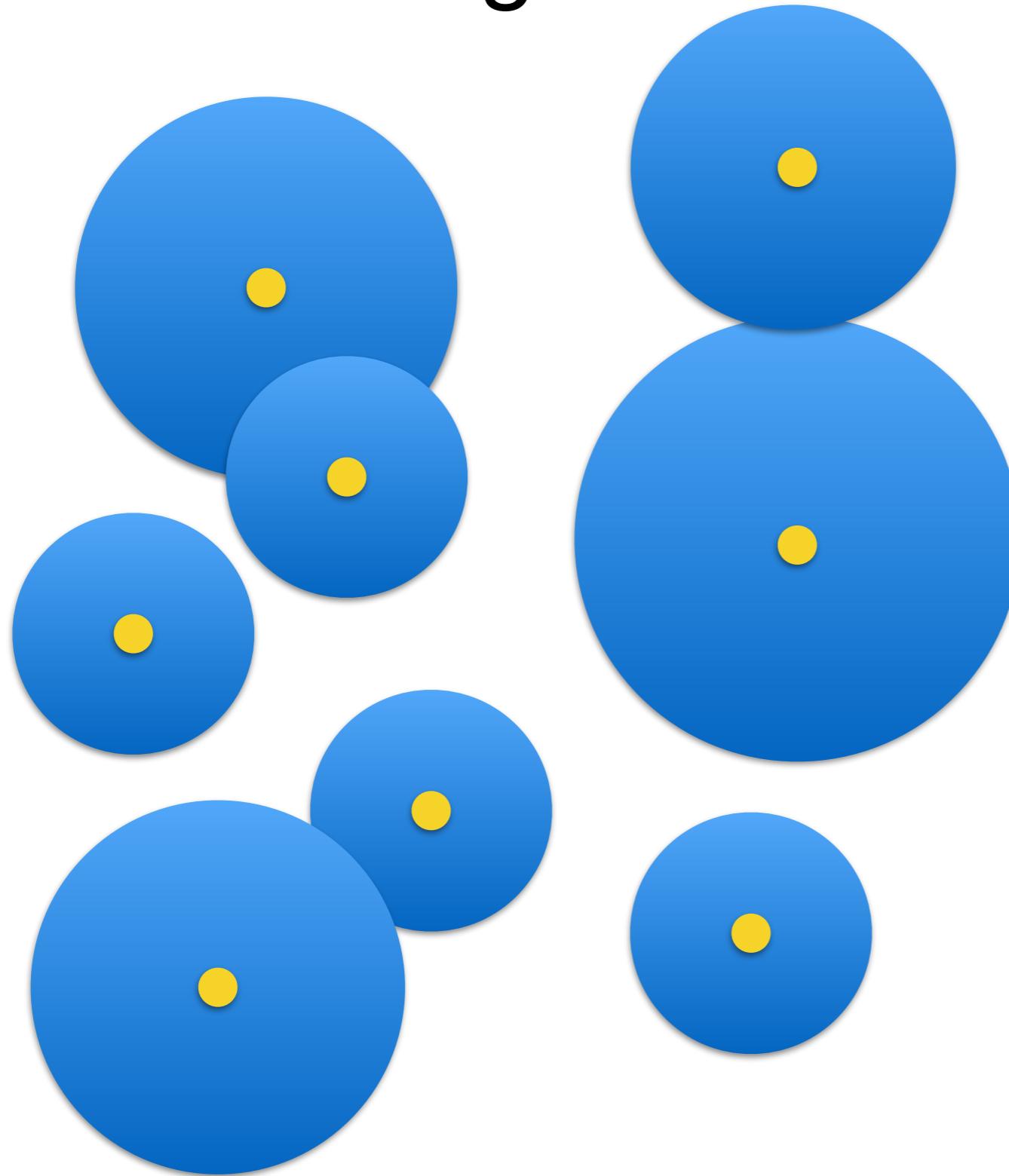


$z \sim 12$

$$\frac{dQ_{\text{ion}}}{dt} = \frac{\dot{n}_\gamma}{n_{\text{H}}} - \frac{Q_{\text{ion}}}{\bar{t}_{\text{rec}}}$$

$$Q_{\text{ion}} = \frac{V_{\text{ion}}}{V_{\text{H}}}$$

Modeling reionization bubbles

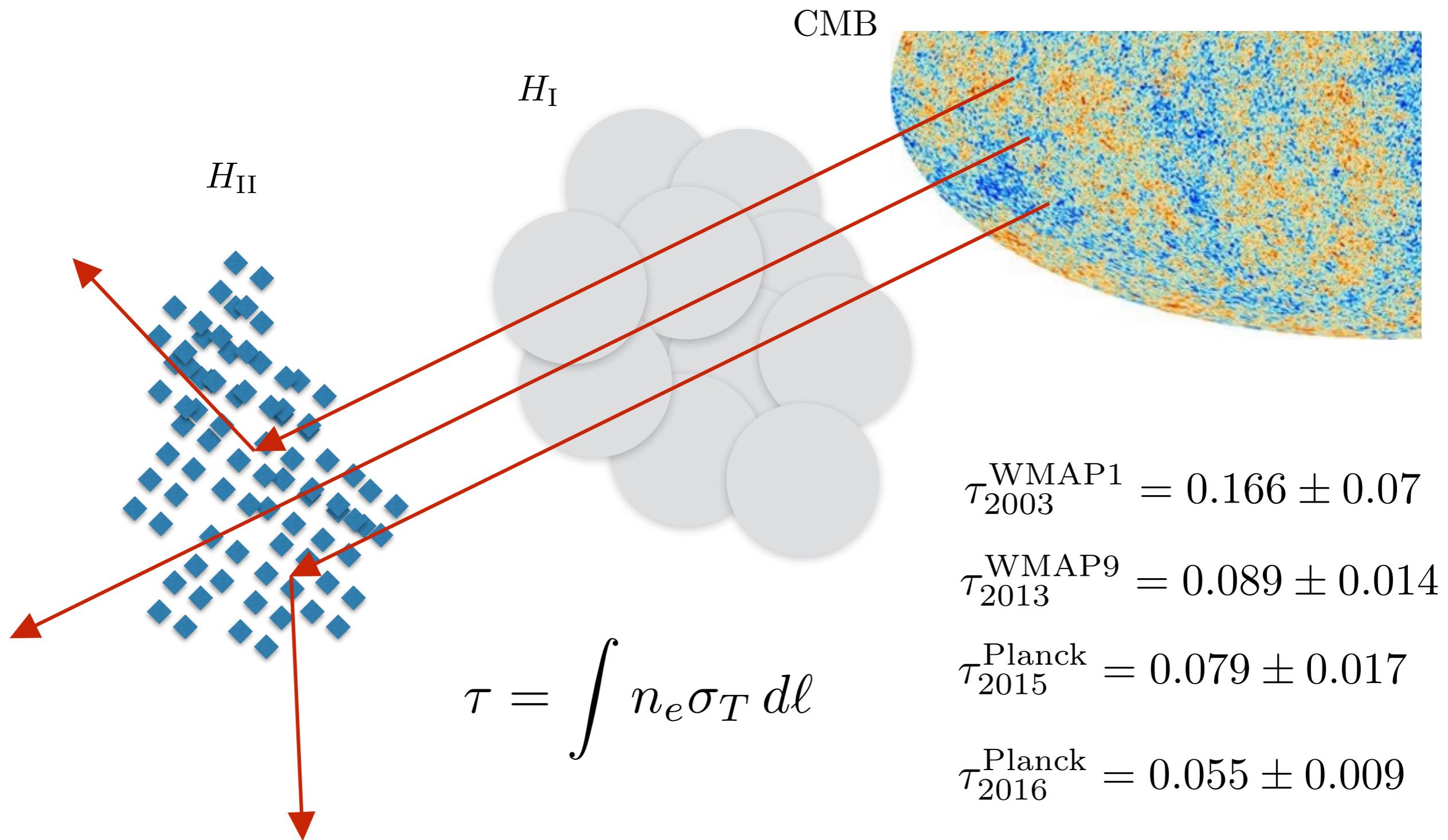


$$\underline{z \sim 6 - 7}$$

$$\frac{dQ_{\text{ion}}}{dt} = \frac{\dot{n}_\gamma}{n_{\text{H}}} - \frac{Q_{\text{ion}}}{\bar{t}_{\text{rec}}}$$

$$Q_{\text{ion}} = \frac{V_{\text{ion}}}{V_{\text{H}}}$$

High redshift CMB constraints



Data

$f_{\text{esc}} < 0.1$ at $z < 3$ (Boutsia, et al., 2011)

$x_{\text{HI}} < 0.16$ at $z < 5.8$ (McGreer, et al., 2014)

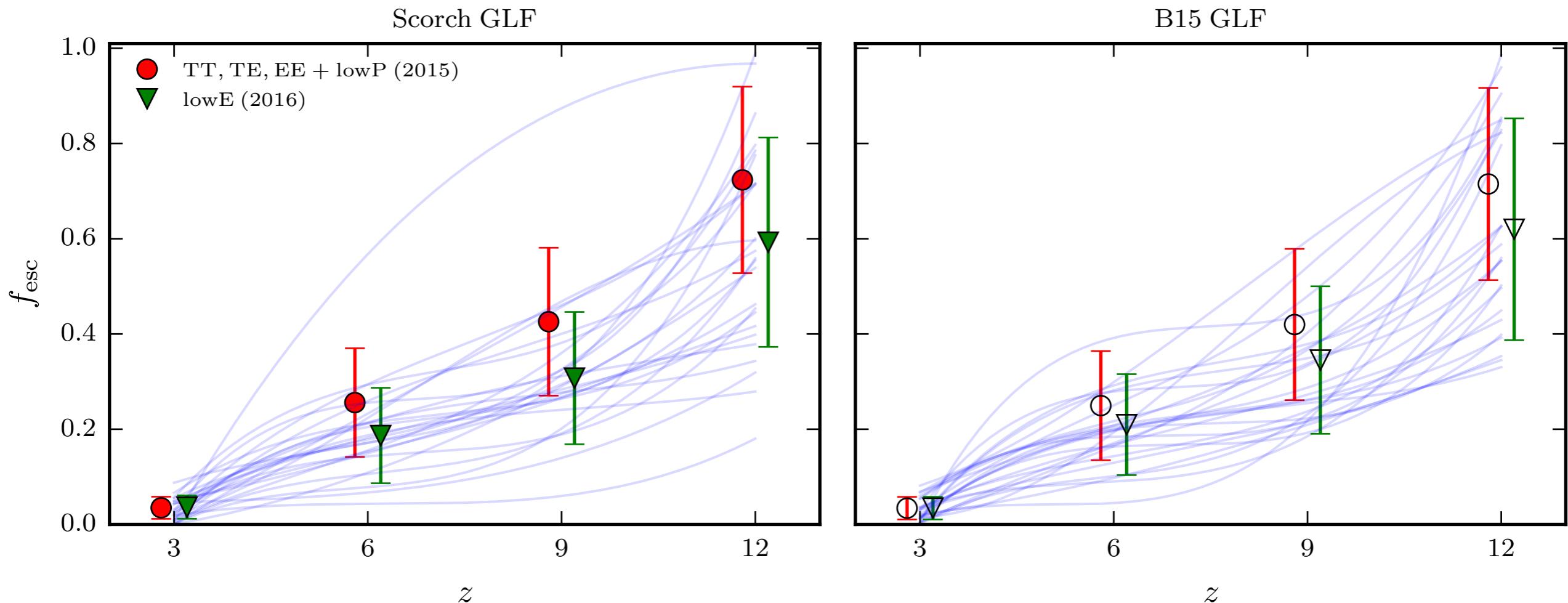
GLF calibrated to HST at $z = 6 - 10$

τ weighted by Planck at $0 \leq z \leq z_{\text{CMB}}$ ($6 \lesssim z \lesssim 12$)

increasing

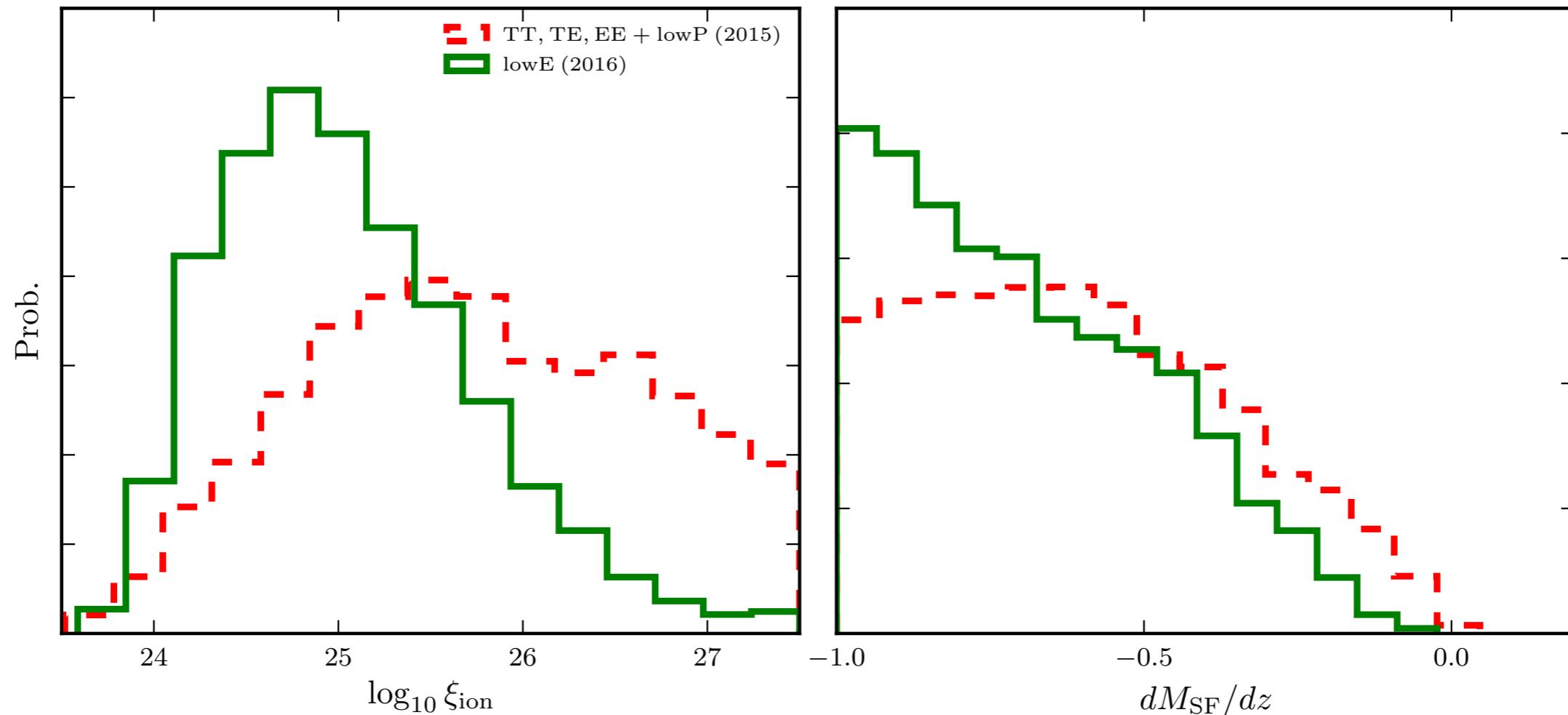
z

Reconstructing the escape fraction



Consistency between escape fraction
reconstructions with different luminosity
functions.

Empirical luminosity function

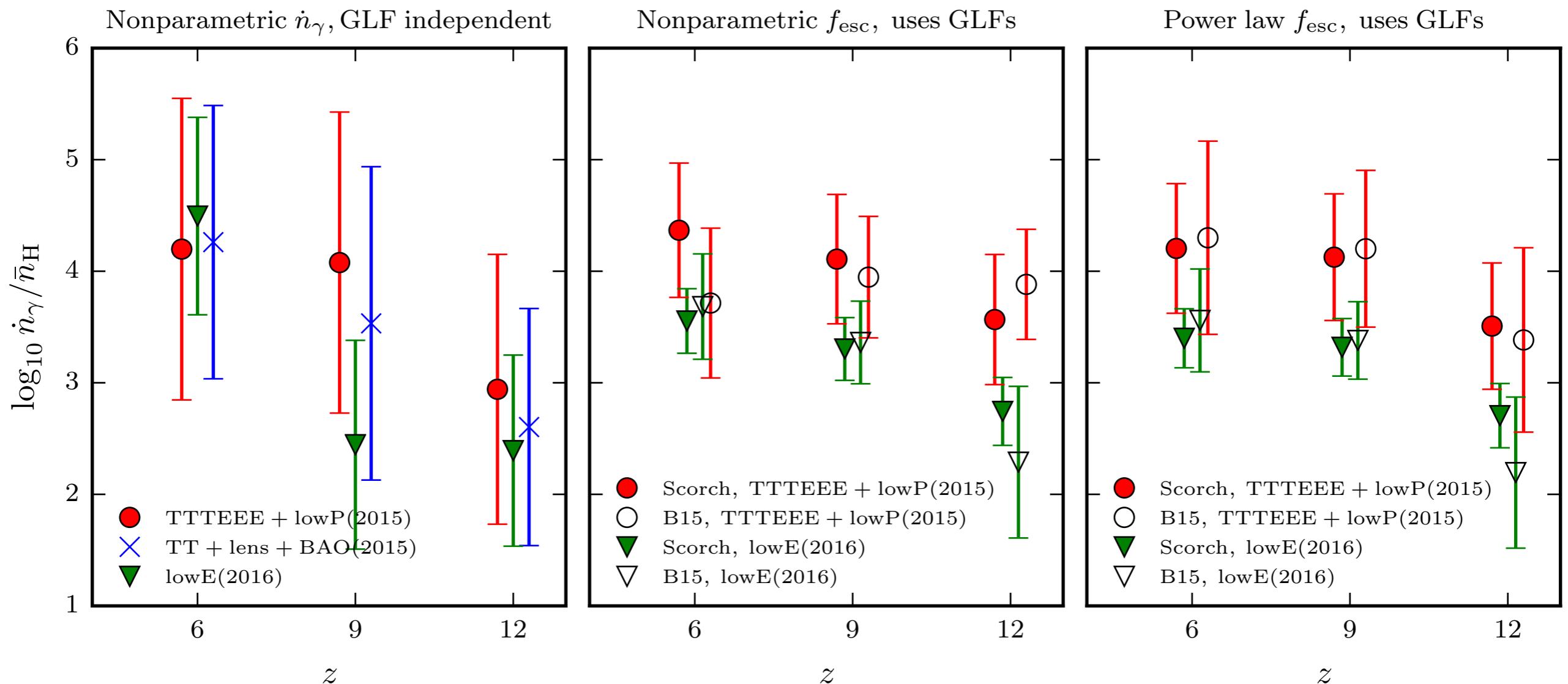


Reduced intrinsic
photon production efficiency
and strong evidence for
evolving limiting magnitude with
latest Planck data

$$L_{\text{ion}} = \xi_{\text{ion}} L_{\text{UV}}$$

$$M_{\text{SF}} = -10 + \frac{dM_{\text{SF}}}{dz}(z - 6)$$

Model-independent photon production rate



Most general non-parametric results show slight preference for suppressed photon production rate at $z \sim 9$.

Conclusion

Performed free-form fit to escape fraction and escaped photon production rate.

Results not strongly dependent on galaxy luminosity function assumptions at high redshift or low luminosity.

Lower Thomson optical depth from Planck 2016 better fit by lowering intrinsic ionizing photon production efficiency (does depend on prior).

Strong evidence for redshift evolution in limiting magnitude of empirical galaxy luminosity functions and escape fraction.