Chaos, Complexity, and Inference (36-462) Lecture 14

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Where Do Heavy Tails Come From?

Change of Variables Some very boring explanations Growing by Multiplying A somewhat boring explanation Critical Fluctuations An exciting and mysterious explanation

More reading: Newman (2005); Mitzenmacher (2004); Schroeder (1991) for some fun examples I can't fit in here

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Change of Variables: Take Logarithms

Suppose $X \sim \text{Pareto}(\alpha, x_{\min})$ Define $Y \equiv \ln X / x_{\min}$

$$F^{\uparrow}(y) = \Pr(Y \ge y) = \Pr(X \ge x_{\min}e^{y})$$
$$= \left(\frac{x_{\min}e^{y}}{x_{\min}}\right)^{-(\alpha-1)} = e^{-(\alpha-1)y}$$
$$Y \sim \operatorname{Exp}(\alpha-1)$$

Conclusion: things only look heavy-tailed because you're measuring the exponential of what you should be measuring Makes sense sometimes... but hard to get behind the idea of "log population" or "log money"

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Change of Variables: Small Denominators

After Sornette (2002) Let $X \sim$ Whatever $Y \equiv X^{-1/a} \Rightarrow$ $f_Y(y) = \alpha \frac{f_X(y^{-\alpha})}{y^{1+\alpha}}$

If $f_X(x) \to c$ as $x \to 0$ then for large y

$$f_Y(y) = O(y^{1+\alpha})$$

Story: we measure the reciprocal of something which should be sensibly-distributed; flat distribution near zero gets turned into a heavy tail towards infinity

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Mixtures of Exponentials

Maguire *et al.* (1952); Beck (2005) Exponential variables, with Γ-distributed rates:

> $X|\Lambda \sim \operatorname{Exp}(\Lambda/s)$ $\Lambda \sim \Gamma(\alpha, 1)$

What is the distribution of X?

$$\Pr(X \ge x) = \int_0^\infty d\lambda \,\lambda^{\alpha-1} \frac{e^{-\lambda}}{\Gamma(\alpha)} \int_x^\infty dy \,\frac{\lambda}{s} e^{-\lambda y/s}$$
$$= \int_0^\infty d\lambda \,\lambda^{\alpha-1} \frac{e^{-\lambda}}{\Gamma(\alpha)} e^{-\lambda x/s}$$
$$= \int_0^\infty d\lambda \,\lambda^{\alpha-1} \frac{e^{-\lambda(1+x/s)}}{\Gamma(\alpha)}$$
$$= \int_0^\infty d\mu (1+x/s)^{-1} \,\mu^{\alpha-1} (1+x/s)^{-(\alpha-1)} \frac{e^{-\mu}}{\Gamma(\alpha)}$$
$$= (1+x/s)^{-\alpha}$$

which is the "Pareto II" distribution

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Multiplicative Growth: Lognormal

Recall central limit theorem: X_i all IID, $\mathbf{E}[X_i] = \mu$, $\operatorname{Var}[X_i] = \sigma^2$, then

$$\sum_{i=1}^{n} X_i \rightsquigarrow \mathcal{N}(n\mu, n\sigma^2)$$

Now let $Y_i = e^{X_i}$:

$$\prod_{i=1}^{n} Y_{i} \rightsquigarrow e^{\mathcal{N}(n\mu, n\sigma^{2})}$$

Issue: CLT is really

$$\frac{1}{n}\sum_{i=1}^n X_i \rightsquigarrow \mathcal{N}(\mu, \sigma^2/n)$$

SO

$$\prod_{i=1}^{n} Y_{i} \rightsquigarrow e^{n \mathcal{N}(\mu, \sigma^{2}/n) + o(n)}$$

and $e^{o(n)}$ is not necessarily small! Put a little differently, the center of the distribution will become log-normal much faster than the tails

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Multiplicative Growth: Exponential Growth with Random Origins

Reed and Hughes (2002) Imagine many piles Each pile grows exponentially

$$X_i(t) = x_0 e^{\lambda(t-T_i)}$$

piles start growing at random times = with constant probability per unit time

$$t - T_i \sim \operatorname{Exp}(\mu)$$

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What is the distribution of pile sizes?

$$\Pr(X_i \ge x) = \Pr\left(e^{\lambda(t-T_i)} \ge x/x_0\right) = \Pr\left(\lambda(t-T_i) \ge \ln x/x_0\right)$$
$$= \Pr\left(t - T_i \ge \frac{\ln x/x_0}{\lambda}\right)$$
$$= e^{-\mu \frac{\ln x/x_0}{\lambda}} = \left(\frac{x}{x_0}\right)^{-\mu/\lambda}$$
$$X \sim \operatorname{Pareto}(\mu/\lambda + 1, x_0)$$

Still works if it's only average size that grows exponentially

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Change of Variables
Multiplicative Growth
Critical Fluctuations
References
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Imagine doing this for cities; $\lambda = \mu = 1/100$

```
> t.start = 2000-rexp(500, rate=1/100)
> summary(t.start)
  Min. 1st Ou. Median Mean 3rd Ou.
                                          Max.
  1327 1852 1927 1892 1969
                                          2000
> plot.new()
> plot.window(xlim=c(min(t.start),2000),ylim=c(1,500),xlab="starting t
> points(t.start,1:500,pch=22)
> axis(3)
> axis(1)
> sizes.now = exp((1/100)*(2000-t.start))
> plot.new()
> plot.window(xlim=c(0, max(sizes.now)), ylim=c(1, 500))
> lines(sizes.now,1:500)
> axis(1)
> axis(3)
> plot.survival.loglog(sizes.now,xlab="present size",ylab="survival fu
> curve(ppareto(x,1,2,lower.tail=FALSE),col="blue",add=TRUE)
```

For US, min $T_i = 1327$ is not crazy, oldest city is Acoma, N.M., from 12th century — but otherwise?





starting times

sizes at *t* = 2000

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Problems: Acoma is not the largest city in the US; largest city is 8×10^6 larger than smallest, not 8×10^2 larger Another model: make X(t) log-normal

$$\ln X(t)/x_0 \sim \mathcal{N}((\mu - \frac{\sigma^2}{2})(t - t_i), \sigma^2(t - t_i))$$

Then $\mathbf{E}[X(t)] = x_0 e^{\mu(t-t_i)}$ This comes from a simple multiplicative growth model, geometric Brownian motion

$$\frac{dX}{dt} = \mu X + \sigma X \xi$$

with ξ = white noise

Unfortunately a real explanation needs stochastic calculus Set $\mu = \sigma^2 = 0.01$

> sizes.gbm = rlnorm(500,(0.01-0.005)*(2000-t.start),0.01*(2000-t.star

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 $\max X / \min X$ now 2×10^6 , a bit small but in the right ballpark



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Yule-Simon Mechanism

Simon (1955); Ijiri and Simon (1977) a.k.a. "the rich get richer", "Matthew Effect", "preferential attachment" . . .

Again with the piles, but now discrete

One lump arrives each time step

Starts new pile with probability ρ

Otherwise joins an existing pile, probability of joining some pile of size *k* is $\propto k$

not necessarily equally likely to join *every* pile of the same size What is the limiting distribution of pile sizes? $N_k(t) =$ number of piles of size k, after t time-steps Assume $N_k(t) \rightarrow \rho_k t$

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If $k \ge 2$,

$$\Pr(N_k(t+1) = N_k(t) + 1) = (1 - \rho)(k - 1)\frac{N_{k-1}(t)}{t}$$
$$\Pr(N_k(t+1) = N_k(t) - 1) = (1 - \rho)k\frac{N_k(t)}{t}$$
$$\mathbf{E}[N_k(t+1)] - N_k(t) = (1 - \rho)\frac{(k - 1)N_{k-1}(t) - kN_k(t)}{t}$$

As $t \to \infty$, we want $N_k(t) \to p_k t$

$$p_{k}(t+1) - p_{k}t = (1-\rho)\frac{(k-1)p_{k-1}t - kp_{k}t}{t}$$

$$p_{k} = (1-\rho)((k-1)p_{k-1} - kp_{k})$$

$$p_{k}(1+(1-\rho)k) = (1-\rho)(k-1)p_{k-1}$$

$$\frac{p_{k}}{p_{k-1}} = \frac{(1-\rho)(k-1)}{1+(1-\rho)k}$$

Define
$$\alpha = 1/(1 - \rho)$$

$$\frac{p_k}{p_{k-1}} = \frac{k-1}{\alpha+k}$$

$$p_k = \frac{k-1}{\alpha+k}p_{k-1}$$

$$= \frac{k-1}{\alpha+k}\frac{k-2}{\alpha+k-1}p_{k-2}$$

$$= \frac{(k-1)(k-2)\dots 2\cdot 1}{(\alpha+k)(\alpha+k-1)\cdot(\alpha+1)}p_1$$

$$= \frac{\Gamma(k)\Gamma(\alpha+1)}{\Gamma(k+\alpha+1)}p_1 = B(k,\alpha+1)p_1$$

Using normalization,

$$p_k = \alpha B(k, \alpha + 1) = O(k^{-\alpha + 1})$$

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Why physicists expect Gaussian fluctuations around equilibrium

Probability of macroscopic variables *M* having value *m* (Einstein fluctuation formula):

$$\Pr(M = m) \propto e^{S(m)}$$

Equilibrium m^* = state of *maximum* entropy, so $\partial S/\partial m = 0$ at m^* ; Taylor expansion in the exponent:

$$\Pr(\mathbf{M} = \mathbf{m}^*) \propto \mathbf{e}^{\mathbf{S}(\mathbf{m}^*) + \frac{1}{2} \frac{\partial^2 S(\mathbf{m}^*)}{\partial \mathbf{m}^2} (\mathbf{m} - \mathbf{m}^*)^2 + \text{h.o.t.}}}$$
$$\propto \mathbf{e}^{\frac{1}{2} \frac{\partial^2 S(\mathbf{m}^*)}{\partial \mathbf{m}^2} (\mathbf{m} - \mathbf{m}^*)^2 + \text{h.o.t.}}$$

drop the h.o.t.

$$M \sim \mathcal{N}(m^*, -\frac{\partial^2 S(m^*)}{\partial m^2})$$

What's really going on

correlations are short range

- \Rightarrow rapid approach to independence, exponential mixing
- \Rightarrow central limit theorem for averages over space (and time)
- \Rightarrow Gaussians

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Phase Transitions

See Yeomans (1992) for nice introduction Basically, bifurcations: behavior changes suddenly as temperature (or pressure or other control variable) crosses some threshold First order: entropy is discontinuous at critical point Examples: ice/water at 273K (and standard pressure); water/steam at 373K order parameter is discontinuous Second order: *derivative* of entropy is discontinuous Example: "Curie point", permanent magnetization/not in iron 1043K order parameter continuous but with sharp kink like amplitude of limit cycle in period-doubling Focus on continuous (second-order) case

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Critical fluctuations

Entropy story breaks down because derivatives $\rightarrow \pm \infty$

Competition between two phases, no preference, one can pop up in the middle of the other

Fluctuations get arbitrarily large \Rightarrow long-range correlations \Rightarrow slow mixing (if any)

Assemblage becomes self-similar: magnify a small part and it looks just like the whole thing ("renormalization")

only strictly true for infinitely big assemblages

averaging must lead to a self-similar distribution

Power laws are self-similar (scale-free)

Conclusion: at critical point, expect to see power law distributions

Landau and Lifshitz (1980); Keizer (1987) are good on details but advanced

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Theory of phase transitions / critical phenomena / order parameters / renormalization one of the key developments in physics over the last half century (Yeomans, 1992; Domb, 1996)

 \Rightarrow physicists think criticality is Very Cool Criticality \Rightarrow power law distributions *so* physicists tend to think:

(ii) is called "the fallacy of affirming the consequent"

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Self-Organized Criticality

See Miller and Page; papers: Bak *et al.* (1987); Carlson and Swindle (1995); Dickman *et al.* (2000); Bak (1996) if *heavily* salted No externally set control parameter Instead, external driving + interactions tend to keep the system towards a critical point Turns out (Dickman *et al.*, 2000) that this is another version of the same story, only with the driving rate tuned very low

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Morals

- There are many ways to obtain heavy-tailed distributions, with or without power law tails
- Some of these mechanisms make different predictions about the distributions
- Even if they do not, they make different predictions about the dynamics
- Both distributions and dynamics can be used to learn about mechanisms

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